

SUSY Phenomenology & Experimental searches

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Slides available at:

<http://www.hep.ph.ic.ac.uk/~tapper/lecture.html>

Objectives

- Know what Supersymmetry (SUSY) is
- Understand qualitatively the problems with the Standard Model that SUSY may be able to solve
- Know somewhat quantitatively the bounds that the different motives for SUSY impose on its mass scale
- Be able to understand limits on SUSY presented in a variety of models
- Understand in general terms search strategy for direct detection at collider experiments
- Know where to find more detailed information

Outline

- What's wrong with the Standard Model?
- What's SUSY and why does it help?
- The usual models
- How do we search for SUSY (at colliders)?

Not included

- Nothing about the Higgs
 - See Gavin Davies's lectures on the Higgs
- Nothing about connection to dark matter experiments
 - Direct searches covered by Henrique et al.
- Combined fits to collider/cosmo/low energy data
 - See Rob Bainbridge's lectures

Reading list

- Vast literature on Supersymmetry
- Three of my favourites:
 - A Supersymmetry Primer (S. Martin)
 - <http://arxiv.org/abs/hep-ph/9709356>
 - Supersymmetry facing Experiment (Pape & Treille)
 - <http://iopscience.iop.org/0034-4885/69/11/R01>
 - SUSY and Such (S. Dawson)
 - <http://arxiv.org/abs/hep-ph/9612229>
- Check any day on hep-ph to see latest papers
 - <http://arxiv.org/archive/hep-ph>

First a poll...

- Who “believes” in supersymmetry?

(even if you don't completely understand what it is)

- Why study this?

What's wrong with the SM?

- Why three generations?
- Neutrinos have mass
- Nineteen free parameters
- Doesn't include gravity or dark matter/energy
- Hierarchy problem
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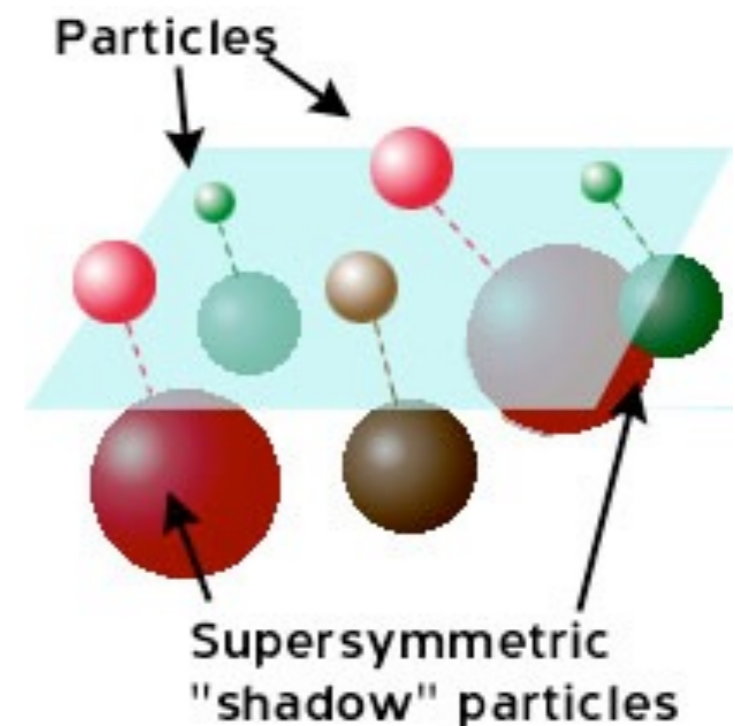
What is supersymmetry?

- If you want to add a symmetry to the Standard Model the Coleman-Mandula theorem says a symmetry connecting fermions and bosons is the only extra (not in the Poincare group) external symmetry you may add and preserve interactions

- Postulate a symmetry such that:

$$\hat{O}|f\rangle = |b\rangle \quad \hat{O}|b\rangle = |f\rangle$$

- Where f (b) are fermions (bosons)
- Particle number is conserved



What is supersymmetry?

- In the Standard Model particles are arranged into multiplets, e.g.

$$\begin{pmatrix} e_L \\ \nu_e \end{pmatrix} \begin{pmatrix} u_L \\ d_L \end{pmatrix}$$

- Supersymmetry follows the same idea:
 - “Supermultiplets”
 - Same number of bosonic and fermionic degrees of freedom
 - Chiral supermultiplet: SM + SUSY fermions
 - Gauge supermultiplet: SM + SUSY bosons

R-Parity

- Supersymmetry introduces a new multiplicative quantum number named R-parity (R_P)
 - $R_P = (-1)^{3(B-L)+2S}$
 - B, L and S are Baryon no., Lepton no. and Spin
 - $R_P = 1$ SM particles, $R_P = -1$ SUSY particles
- Think of R_P as “superness”
- SUSY theories can conserve or violate R_P
 - Big consequences in phenomenology

Motivations for supersymmetry

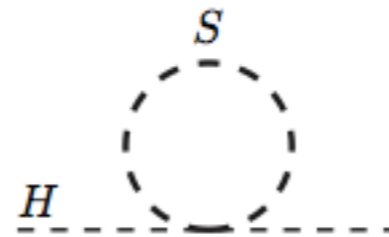
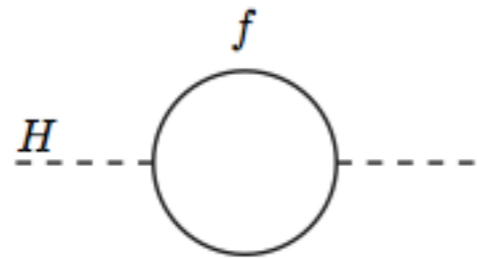
- Naturalness (the hierarchy problem)
- Unification of the EM, weak and strong forces (maybe...)
- Dark Matter candidate

Hierarchy problem

- In the Standard Model there's a Higgs boson
 - Weak scale set by vev of Higgs ~ 246 GeV
 - In principle sets scale for all masses in the theory including the Higgs (now know 125 GeV)
- Gravity scale set by the Planck mass 10^{19} GeV
- Why is the weak scale so different from the scale of gravity?

Hierarchy problem

- Looking in more detail at the Higgs Mass



- Receives corrections from all particles
 - From fermions: $\Delta m_H^2 = \frac{\lambda_f^2}{8\pi^2} \left[-\Lambda^2 + 6m_f^2 \ln \frac{\Lambda}{m_f} \right]$
 - From scalars: $\Delta m_H^2 = \frac{\lambda_S}{16\pi^2} \left[\Lambda^2 - 2m_S^2 \ln \frac{\Lambda}{m_S} \right]$
 - Λ is cut-off scale to protect against infinite corrections
- Quadratic divergence
 - Natural scale for Higgs mass is cut-off scale
 - Can renormalise away the divergence if it has no physical meaning → but what about gravity?

Hierarchy problem

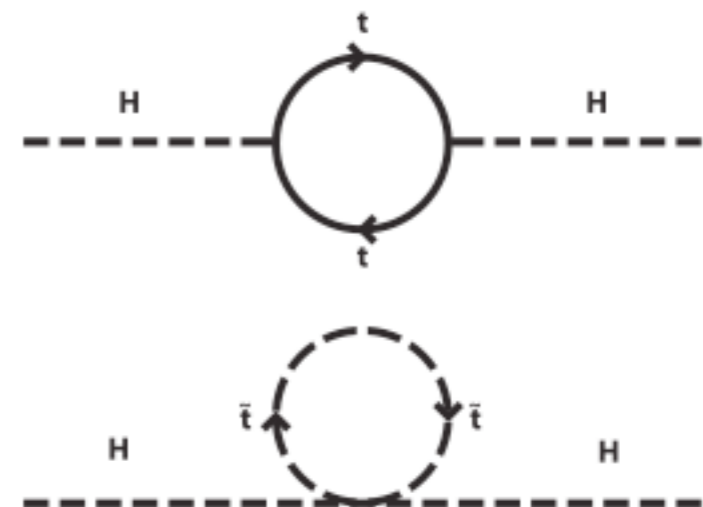
- So the Higgs mass is extremely sensitive to the heaviest particle to which it couples and it's most natural value is close to the largest mass in the theory
- There was a hint to the solution on the previous page:
 - Fermions $\Delta m_H^2 = \frac{\lambda_f^2}{8\pi^2} [-\Lambda^2 + 6m_f^2 \ln \frac{\Lambda}{m_f}]$
 - Scalars $\Delta m_H^2 = \frac{\lambda_S}{16\pi^2} [\Lambda^2 - 2m_S^2 \ln \frac{\Lambda}{m_S}]$
- So if every fermion is accompanied by two scalars with couplings $\lambda_S = \lambda_f^2$ the quadratic divergences cancel
- Relation between couplings by imposing a symmetry
 - Supersymmetry

Hierarchy problem

- The correction then reduces to:

$$\Delta m_H^2 \approx \frac{\lambda_f^2}{4\pi^2} (m_S^2 - m_f^2) \ln \frac{\Lambda}{m_S}$$

- Now divergence is logarithmic (normal in SM)
- Masses can be large but if almost degenerate correction is well behaved
- In Supersymmetry this is basically the top squark cancelling the effect of the top quark
- We'll come back to this when we talk about experimental searches



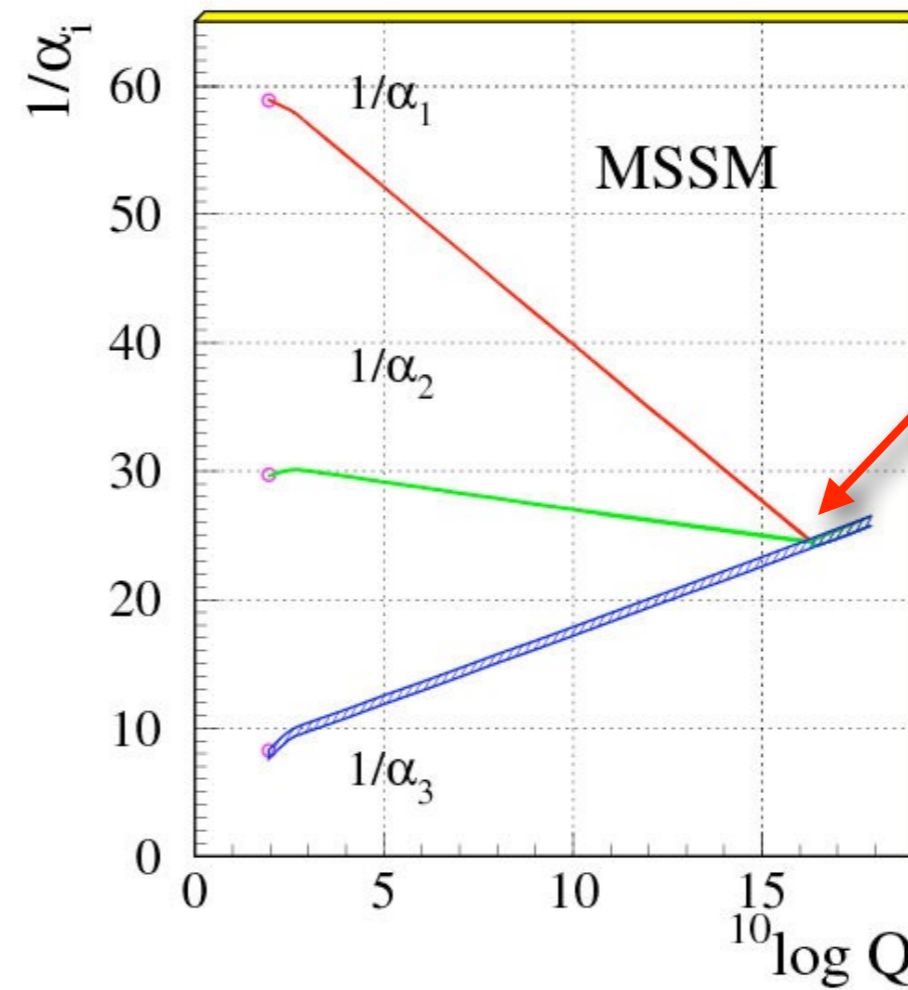
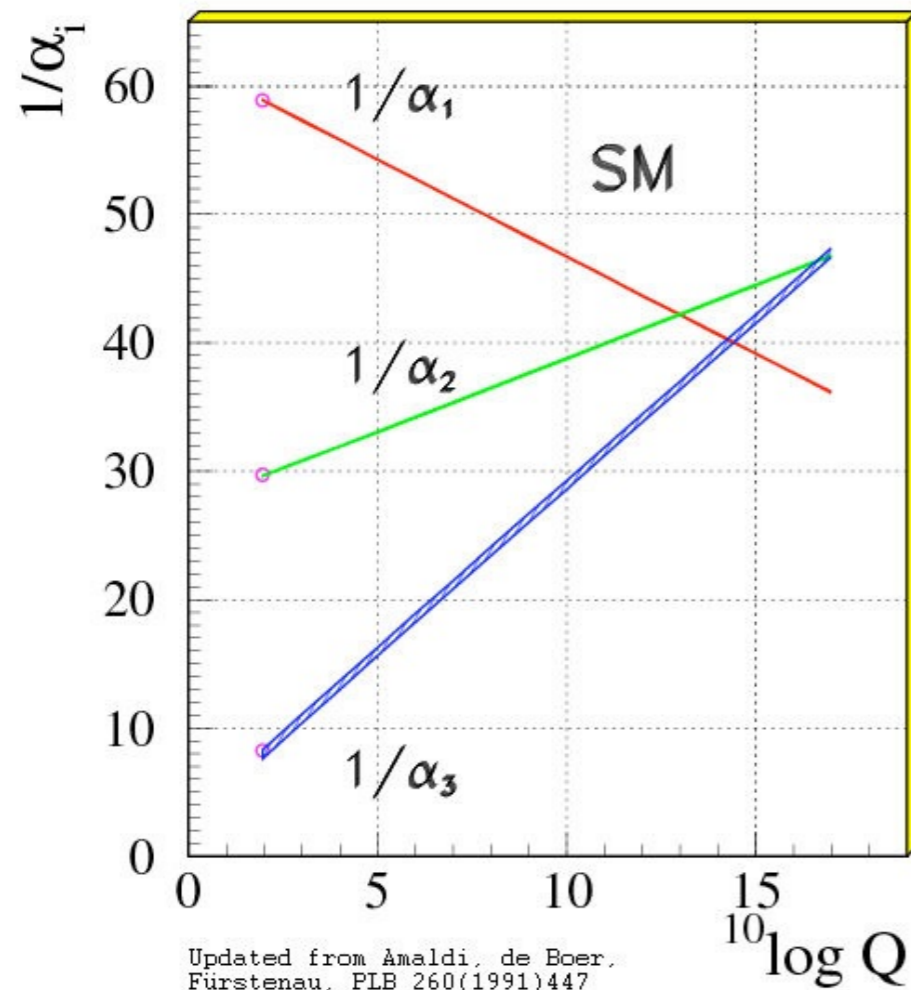
Unification

- A Grand Unified Theory (GUT) unifies the three gauge interactions of the Standard Model
 - Electromagnetic
 - Weak
 - Strong
- Unifies means they are characterised by one gauge symmetry and one coupling constant
- A Theory Of Everything (TOE) would also include gravity

Unification

- Standard Model
 - $SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$
- Theory to unify gauge interactions should include these groups
 - $SU(5), SO(10), \dots$
- Predictions
 - Proton decay, magnetic monopoles,

Unification

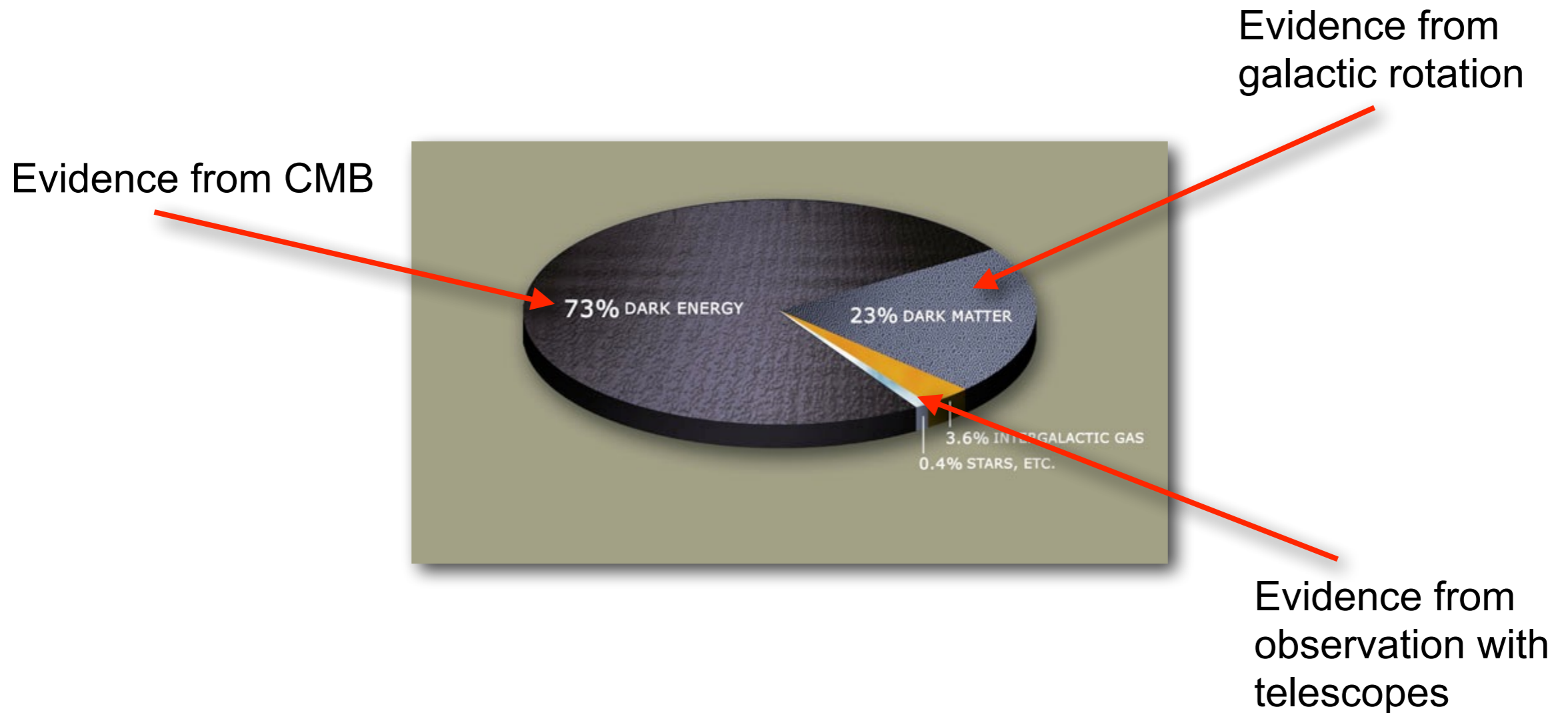


Consistent with proton decay limits

- Couplings run like: $\frac{d}{d(\ln Q)} (\alpha_i^{-1}) = \frac{b_i}{2\pi}$
- Take measured couplings at Z pole and b_i from MSSM
- Unification at 1% level at $\Lambda_{\text{GUT}} \approx 10^{16}$ GeV for masses < 10 TeV

Dark Matter

- **23%** of Universe!



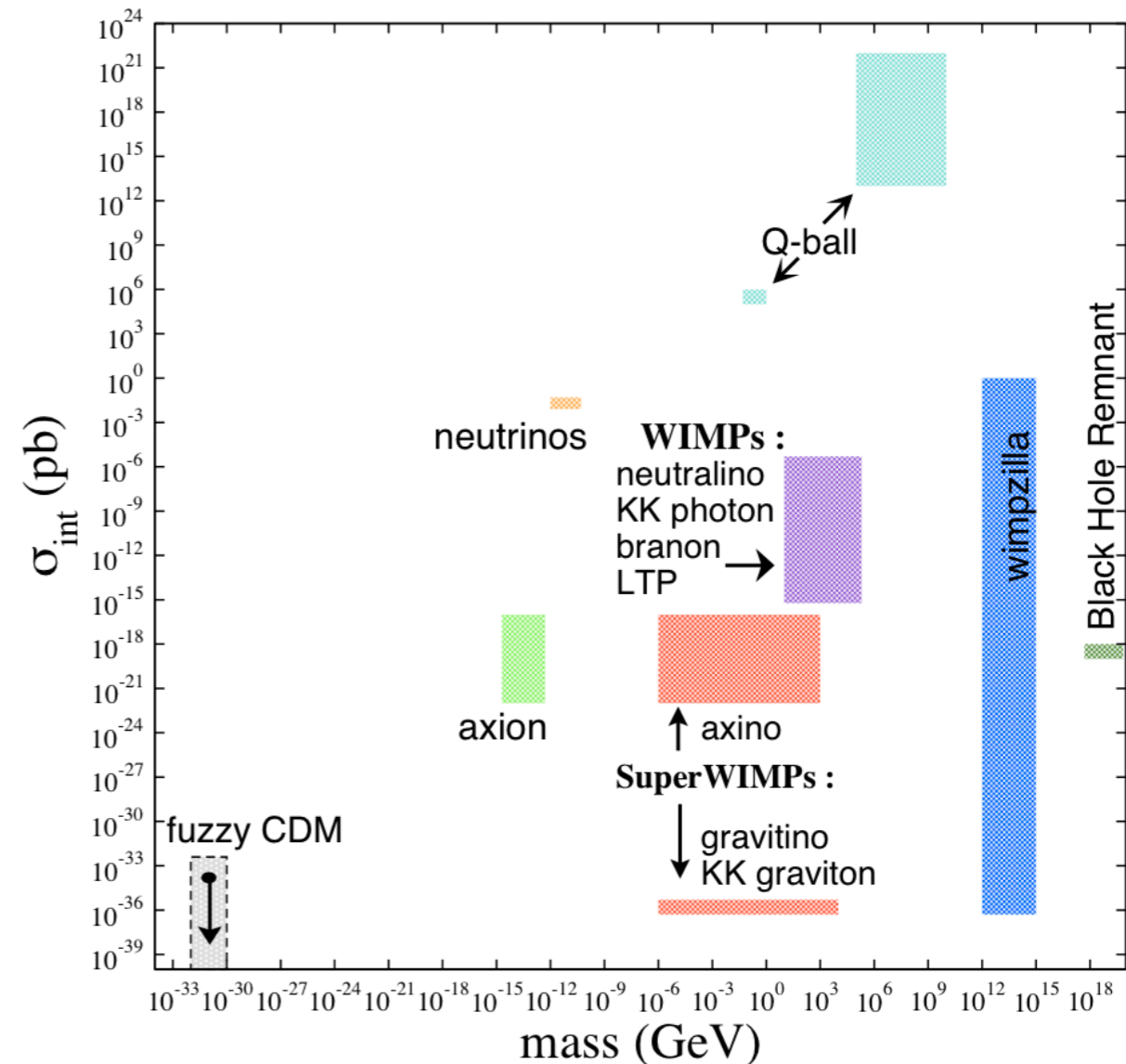
- Many direct searches
- Many different particles proposed →

Dark Matter

- SUSY can provide a dark matter candidate
- No strong or electromagnetic interactions
- WIMP
- Non-baryonic Cold Dark Matter (CDM)

L. Roszkowski, Pramama, 62 (2004) 389

Some Dark Matter Candidate Particles



Dark Matter

- Relies on R-Parity conservation (RPC)
- Remember if R_P is conserved then SUSY particles can only be produced in pairs of a SUSY particle and its antiparticle
- SUSY particles cannot decay directly to SM particles so the lightest SUSY particle has nothing to decay to → stable, weakly interacting Dark Matter candidate
- Can be lightest neutralino, sneutrino, Gravitino....
 - Mass depends on details but for Wino $M_{LSP} < 3 \text{ TeV}$

The MSSM

- **Minimal** Supersymmetric Standard Model
- The least number of particles you can add to the Standard Model to make a viable SUSY model (N=1 supersymmetry)
- Assume R-Parity is conserved (stable proton)
- Each SM particle has a SUSY partner
 - In total $O(120)$ free parameters mostly to do with symmetry breaking →

MSSM Higgs

- Supersymmetry requires two Higgs doublets
 - To cancel gauge anomalies and provide mass to both up and down-type particles
- Since we are not swamped in SUSY particles we assume SUSY is a broken symmetry
 - Aside: could any of the particles we know of be super-partners?
- Many different theories for SUSY breaking
 - Generally spontaneous symmetry breaking in a hidden sector is communicated to the visible sector through corrections to the masses
 - Radiative electroweak symmetry breaking → “more natural”

MSSM particles

Names	Spin	P_R	Gauge Eigenstates	Mass Eigenstates
Higgs bosons	0	+1	$H_u^0 H_d^0 H_u^+ H_d^-$	$h^0 H^0 A^0 H^\pm$
squarks	0	-1	$\tilde{u}_L \tilde{u}_R \tilde{d}_L \tilde{d}_R$	(same)
			$\tilde{s}_L \tilde{s}_R \tilde{c}_L \tilde{c}_R$	(same)
			$\tilde{t}_L \tilde{t}_R \tilde{b}_L \tilde{b}_R$	$\tilde{t}_1 \tilde{t}_2 \tilde{b}_1 \tilde{b}_2$
sleptons	0	-1	$\tilde{e}_L \tilde{e}_R \tilde{\nu}_e$	(same)
			$\tilde{\mu}_L \tilde{\mu}_R \tilde{\nu}_\mu$	(same)
			$\tilde{\tau}_L \tilde{\tau}_R \tilde{\nu}_\tau$	$\tilde{\tau}_1 \tilde{\tau}_2 \tilde{\nu}_\tau$
neutralinos	1/2	-1	$\tilde{B}^0 \tilde{W}^0 \tilde{H}_u^0 \tilde{H}_d^0$	$\tilde{N}_1 \tilde{N}_2 \tilde{N}_3 \tilde{N}_4$
charginos	1/2	-1	$\tilde{W}^\pm \tilde{H}_u^\pm \tilde{H}_d^\mp$	$\tilde{C}_1^\pm \tilde{C}_2^\pm$
gluino	1/2	-1	\tilde{g}	(same)
goldstino (gravitino)	1/2 (3/2)	-1	\tilde{G}	(same)

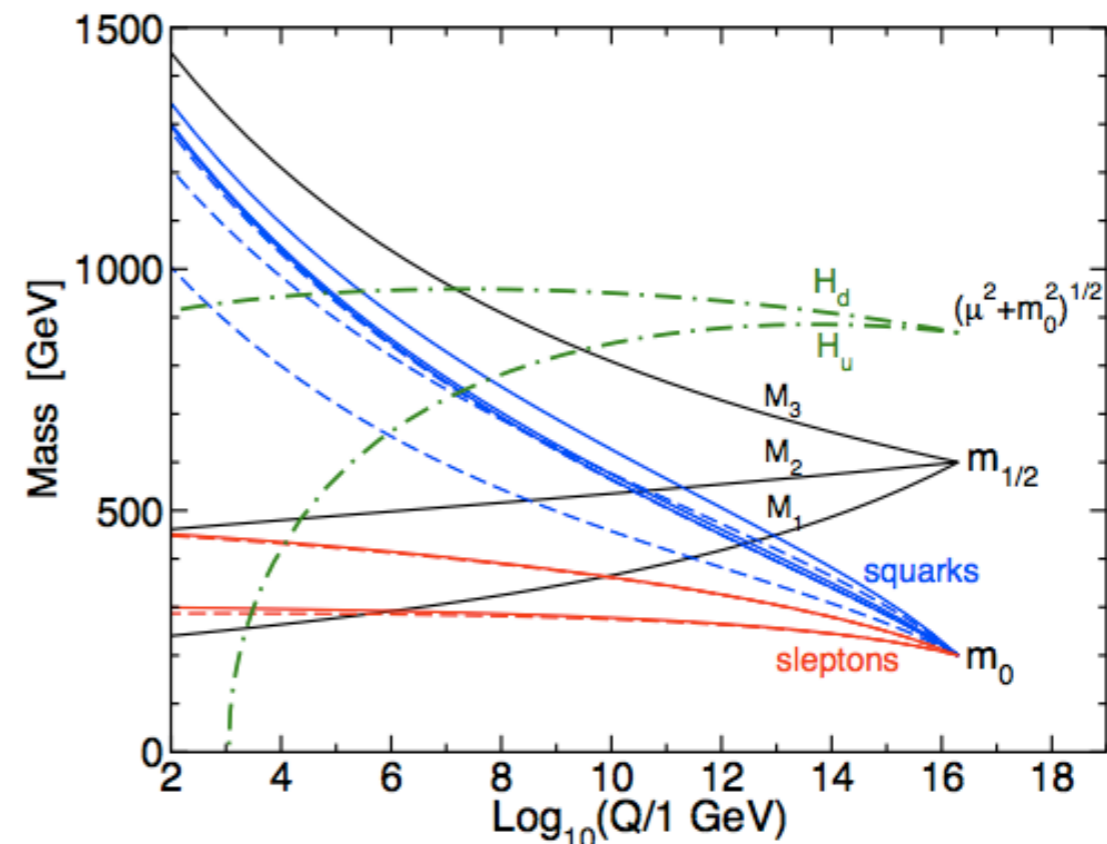
Neutralinos and charginos are often denoted as $X^{0,\pm} \chi$

The CMSSM

- Take a step further in simplification
 - **Constrained** MSSM (sometimes mSUGRA)
- Impose GUT scale (M_{Pl}) relations on the MSSM
 - Set all scalar masses to one value m_0
 - Set all gaugino masses to one value $m_{1/2}$
 - Set trilinear couplings to one value A_0
 - Set ratio of Higgs doublet VEVs to $\tan\beta$
- Gravity-Mediated Supersymmetry Breaking

The CMSSM

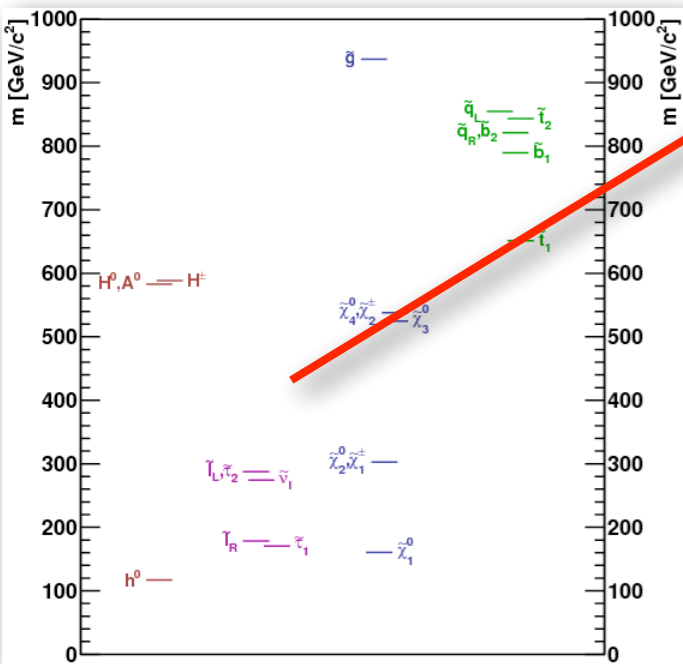
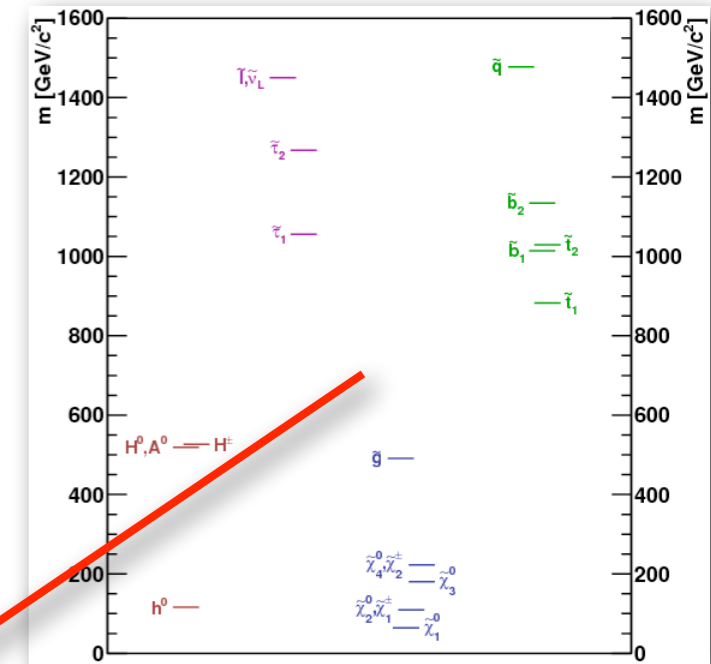
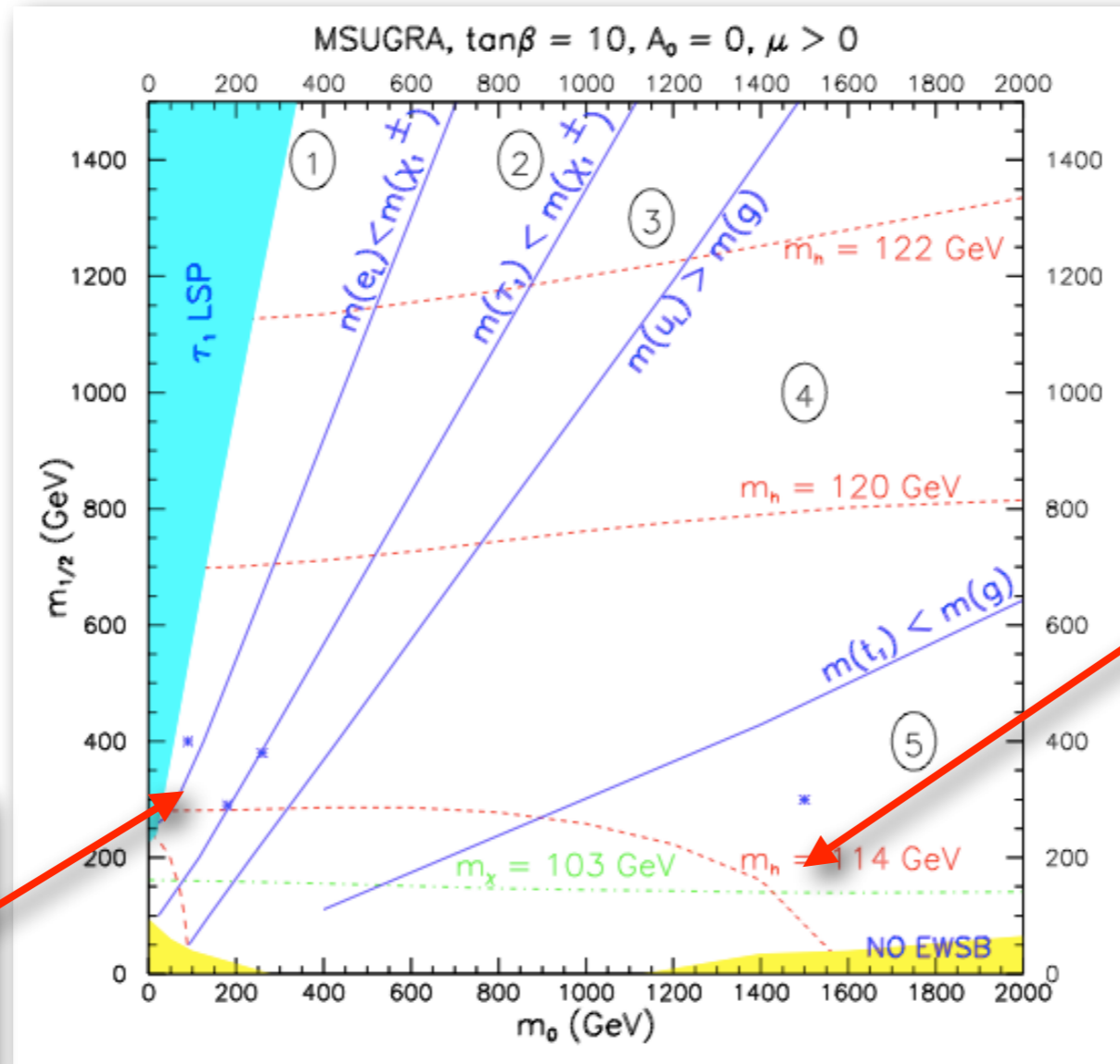
- A full viable SUSY theory in principle (nobody believes this is nature though)
- Very convenient way to simplify SUSY to just 4 parameters (and a sign)
- Widely used to quote results of experimental searches (next lecture)



The CMSSM

- Has a wide range of phenomenologies, but beware!
 - Unifications at the GUT scale lead to fixed ratios between couplings and therefore masses
 - $M_{\text{gluino/squark}} : M_{\text{Chargino}} : M_{\text{Neutralino}} \approx 6 : 4 : 1$
 - Large splittings between gluino, squarks and neutralinos \rightarrow easy signals to observe

The CMSSM



Summary

- Supersymmetry is a theory which postulates a new symmetry between fermions and bosons
- Best studied extension to the Standard Model with vast literature
- Has the potential to solve some of the most serious problems in the Standard Model quite naturally
- Next time: how we search for Supersymmetry at colliders